

Modeling Laser-Plasma Interactions in MagLIF Experiment on NIF

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Modeling Laser-Plasma Interactions in MagLIF Experiment on NIF

Anomalous Absorption Meeting

D. J. Strozzi, R. L. Berger, A. B. Sefkow, S. H. Langer, T. Chapman, B. Pollock, C. Goyon, J. Moody

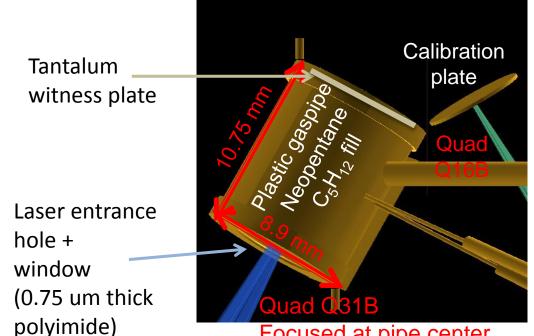
5 May 2016





MagLIF shot on NIF gave excellent laser propagation and good agreement with modeling

MagLIF NIF Target (View from GXD x-ray camera)



This talk: shot N160128 Repeat, better diagnostics: N160425

> B. Pollock, R2-1: Prior talk More on expts

Successfully demonstrated laser propagation at MagLIF fusion-gain scale

Focused at pipe center

Summary: MagLIF NIF shots modeled with radhydro and LPI codes

Modeling tools

- HYDRA: ICF radiation-hydrodynamic code
 - Agrees with laser propagation down tube
 - Provides plasma conditions for LPI modeling
- Gain spectrum: linear gain exponents integrated along laser rays
 - 1D, linear, kinetic, fast no speckles, filamentation, nonlinear kinetics
- pF3D: paraxial envelope propagation code
 - Massively parallel, 3D NIF-relevant volumes [R. Berger, S. Langer Tuesday]

SRS: peak reflectivity ~ 0.3%, from fill gas

- Measured and gain spectra: close, contain two distinct wavelengths
- pF3D: two SRS wavelength groups: dominant one agrees with data

SBS: Peak reflectivity ~ 3% when laser hits Ta plate

- Gain spectrum close to data, but gain from gas not Ta
- pF3D modeling ongoing

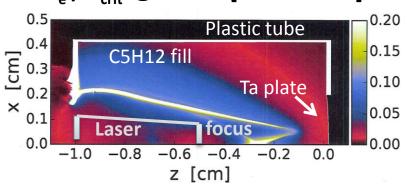


MagLIF NIF shot follows standard NIF "warm" (293 K) surrogacy approach

NIF TARGET

- Gaspipe: 1 cm long, 1 cm diameter
- **Thin window:** 0.75 um polyimide
 - Use same warm and cryo
- MagLIF D₂ fill breaks window @ STP
- Use large hydrocarbon: match n_e
- Fill: neopentane C₅H₁₂ @ 1 atm.
 - $n_e = 0.116 n_{crit}$ fully ionized
 - Same n_e as D₂ at 3.5 mg/cm³
- No imposed B field: 10-20 T in 2017?

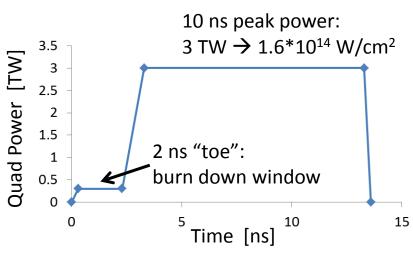
n_e / n_{crit} @ 8.5 ns [HYDRA sim.]



NIF LASER: well-conditioned

- Wavelength: 351 nm "3ω"
- One 30° cone quad (4 beams) Q31B
- Nominal phase plates, F=8 for quad
- "Checkerboard" polarization smoothing
- SSD: 45 GHz
- Focal spot: ellipse, radii (824, 590) um

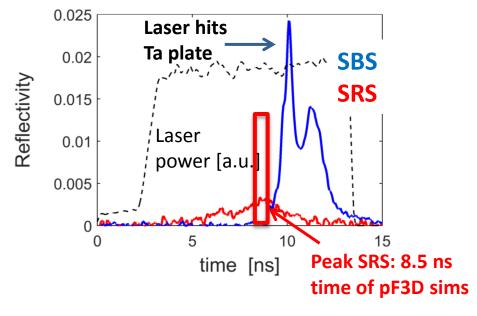
Laser pulse



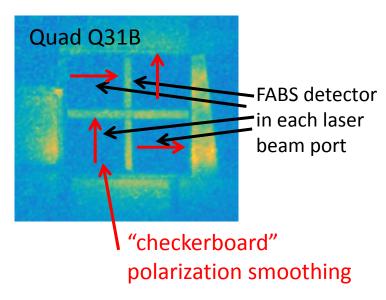


Low power and intensity gave low backscatter, some SBS when laser hits Tantalum plate

Backscatter into FABS detector = lens aperture



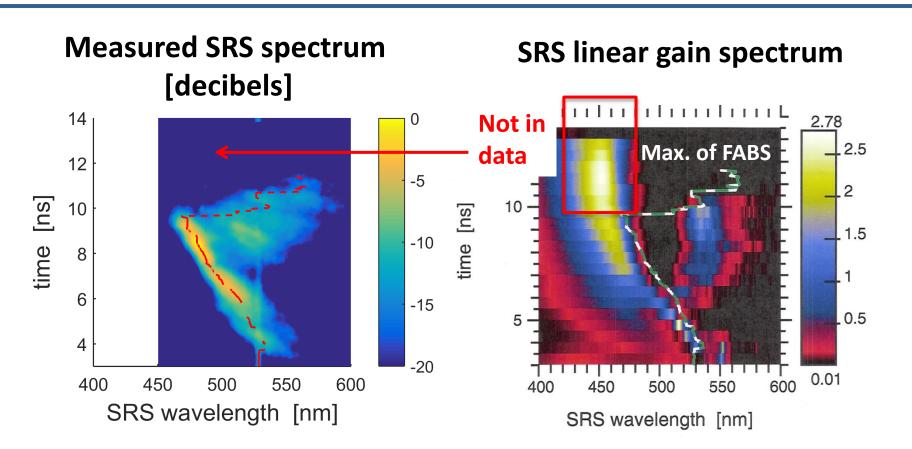
SRS NBI (Near Backscatter Imager) Image centered on laser ports



- Laser hits Ta plate at 10 ns close to x-ray camera data
- Additional backscatter on NBI plate outside of lens ~ few *FABS: analysis ongoing

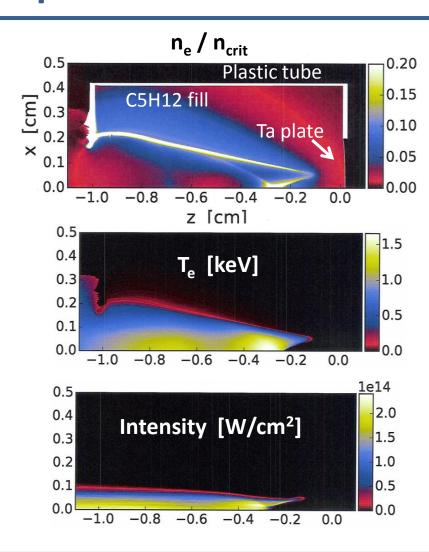


SRS data and gain spectrum qualitatively similar before 10 ns



- Main feature moves to shorter wavelength with time \rightarrow lower n_e
- Longer wavelength feature appears late in time

Plasma conditions from HYDRA run at 8.5 ns: peak measured SRS

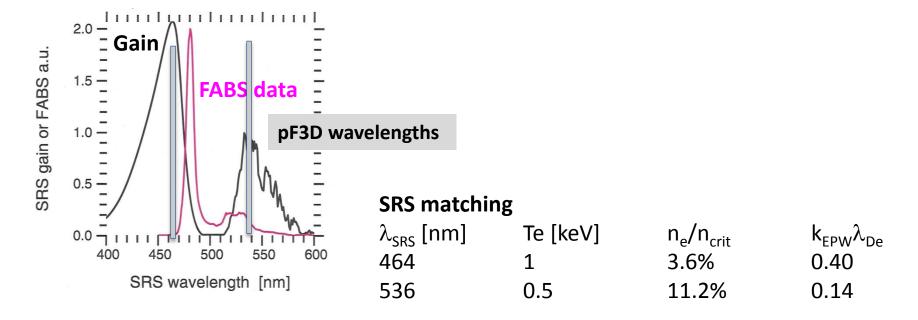


HYDRA run:

- No MHD
- f=0.05 electron heat flux limit
- DCA non-LTE atomic physics



SRS at 8.5 ns: two features in data and gain spectrum



T_e chosen from pF3D results

pF3D*: paraxial envelope light propagation code, massively parallel

*R. L. Berger, C. H. Still, E. A. Williams, A. B. Langdon, *Phys. Plasmas* 1998

Light wave vector potential:

$$\vec{A}_0(\vec{x},t) = \frac{1}{2} \tilde{A}_0(\vec{x},t) \hat{p} \exp i(-\omega_0 t + \phi_0) + cc$$
Slavely we wind:

Polarization:

Slowly-varying envelope

fixed, in xy plane

Envelopes evolved:

- Laser light
- SRS light 1 or 2 wavelength groups
- SRS Langmuir wave 1 or 2 groups
- SBS light
- SBS ion wave: no time enveloping

Background hydro w/ ponderomotive force:

- Filamentation
- Cross-beam energy transfer

Laser envelope equation:

$$\begin{bmatrix} \partial_t + v_{g0} \partial_z - i \frac{(c^2/\omega_0) \nabla_\perp^2}{1 + \left(1 + k_0^{-2} \nabla_\perp^2\right)^{1/2}} + \nu_0 + i \partial_t \phi_0 + \frac{1}{2} \partial_z v_{g0} \end{bmatrix} \tilde{A}_0 \propto \delta n_{ef} \tilde{A}_0 + \frac{1}{2} \delta n_a \tilde{A}_B + \frac{1}{2} \delta n_l \tilde{A}_R$$
 Advection: Damping DAW phase shift Feit-Fleck form

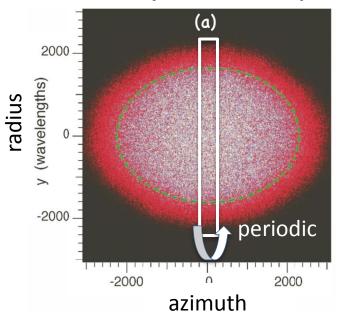
damping limit

pF3D "Letterbox" run for backscatter: routine vs. "heroic" 3D run

"Letterbox": slice in one transverse direction

 Same intensity distribution and speckle statistics as full beam

Laser Intensity in transverse plane



Computing resources

Spatial zoning: $dx = dy = 2 \lambda_0$, $dz = 3 \lambda_0$

Plasma volume 1.9 mm³

Zones: 3.9 billion

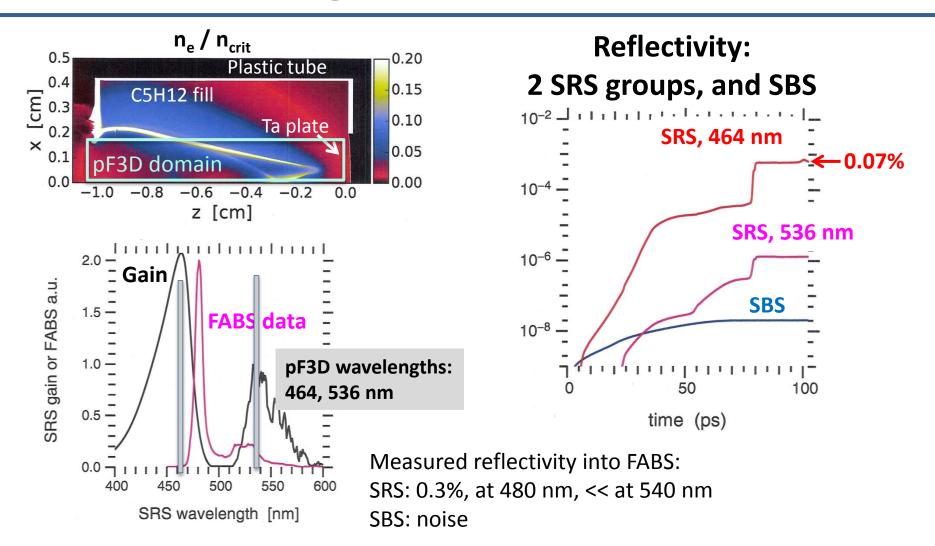
LLNL Sequoia machine: 8192 cpu's, ~ 1 day



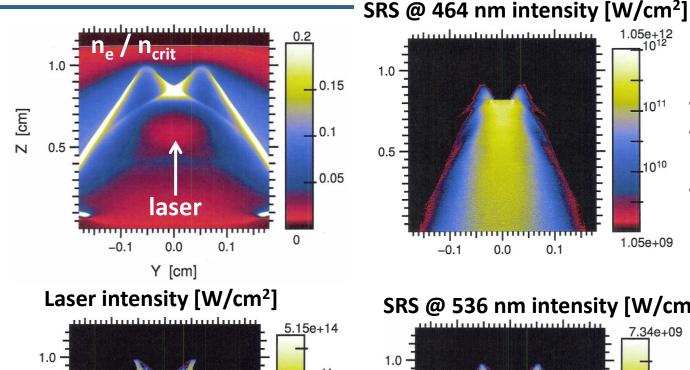


Sample letterbox: D. Hinkel et al., Phys. Plasmas 2008

Peak SRS (8.5 ns): pF3D agrees with data: shorter wavelength SRS dominates, SBS small

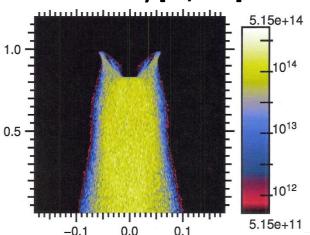


Peak SRS: SRS develops at end of laser path

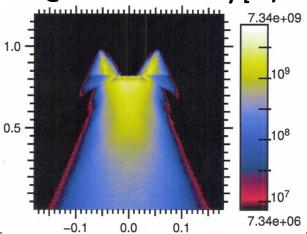


pF3D

- Time 104 ps
- Intensities on different log scales
- Aspect ratio not unity

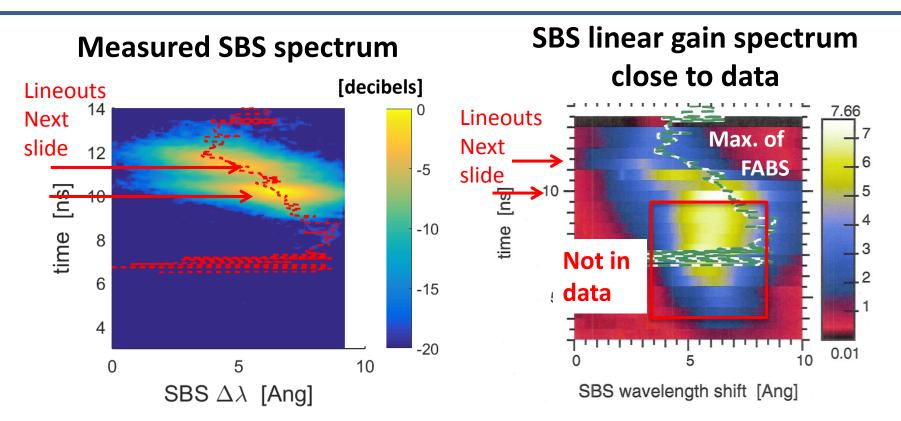


SRS @ 536 nm intensity [W/cm²]



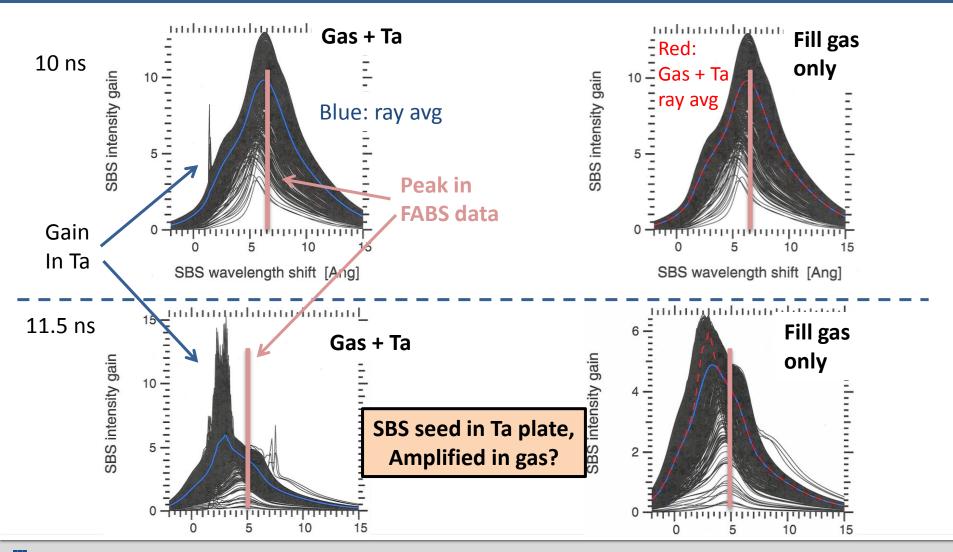


Late-time SBS gain spectrum consistent with data



Late-time SBS occurs when laser hits tantalum back plate: but where is it coming from?

SBS gain spectra late in time: most gain coming from gas, some at short wavelength from Ta



Conclusions and future work

Modeling

- HYDRA correctly gives laser propagation, based on x-ray camera data
- SRS: two wavelengths in gain and data, pF3D gives same dominant one as data
- SBS burst when laser hitting Ta back plate, but gain in gas at that time

Future NIF shots

- Push to higher backscatter risk:
 - Higher intensity
 - Higher fill density
- Cryogenic D₂ fill, thin window: ignition relevant, instead of warm surrogate C₅H₁₂
- Imposed B field: 10-20 T in 2017?

Warm C5H12 fill, no imposed B field:
Successful laser propagation at MagLIF fusion-gain scale

Cryogenic D2 fill, imposed B field:
Will test complete MagLIF scheme – to be done soon...

B. Pollock, R2-1: Prior talk More on expts



BACKUP BELOW

SBS shift in high Z plasma:

$$\delta\lambda[\mathring{\rm A}]\approx 7.3\left(\frac{Z}{A}T_e[keV]\right)^{1/2}\left(1+\frac{\vec{u}\cdot\vec{k}_0}{c_{ac}}\right) \quad \Longrightarrow \quad 4.9\mathring{\rm A} \quad \begin{array}{l} {\rm Tantalum: \ A=181, \ Z=42} \\ {\rm T_e=2 \ keV, \ u=0} \end{array}$$

